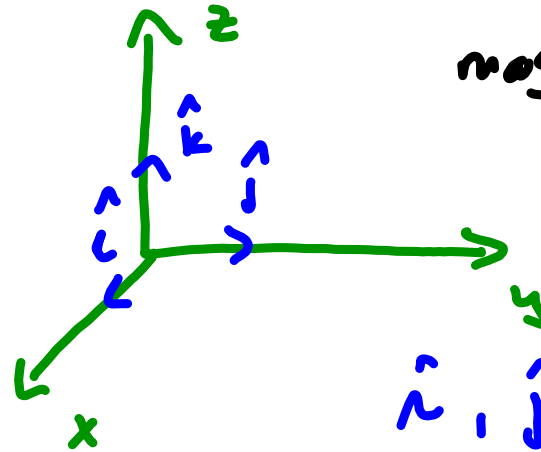


Governing equations

First: math review

$$\vec{V} = u \hat{i} + v \hat{j} + w \hat{k}$$



magnit. $\left\{ \begin{array}{l} u \text{ zonal wind} \\ v \text{ meridional wind} \\ w \text{ vertical wind} \end{array} \right.$

$\hat{i}, \hat{j}, \hat{k}$ direction

$$\vec{V} = u \hat{i} + v \hat{j} + w \hat{k}$$

$$\frac{d\vec{v}}{dt} = \frac{du}{dt} \hat{i} + \frac{dv}{dt} \hat{j} + \frac{dw}{dt} \hat{k} \quad \text{acceleration}$$

if we ignore derivatives
of unit vectors

(acceleration of

We need this for $\sum \vec{F} = m \frac{d\vec{v}}{dt}$ ^{Earth}

Partial derivatives

T temperature

$T(x, y, z, t)$
↑ time
} space

$$u = \frac{dx}{dt}$$

$$v = \frac{dy}{dt}$$

$$w = \frac{dz}{dt}$$

$$\frac{dT}{dt} = \left(\frac{\partial T}{\partial t} \right)_{x,y,z} dt + \left(\frac{\partial T}{\partial x} \right)_{y,z,t} dx + \left(\frac{\partial T}{\partial y} \right)_{x,z,t} dy + \left(\frac{\partial T}{\partial z} \right)_{x,y,t} dz$$

$$\frac{dT}{dt} = \frac{\partial T}{\partial t} + \frac{\partial T}{\partial x} \frac{dx}{dt} + \frac{\partial T}{\partial y} \frac{dy}{dt} + \frac{\partial T}{\partial z} \frac{dz}{dt}$$

$$\textcircled{*} \quad \frac{dT}{dt} = \frac{\partial T}{\partial t} + u \frac{\partial T}{\partial x} + v \frac{\partial T}{\partial y} + w \frac{\partial T}{\partial z}$$

more math overview

$$\vec{\nabla} = \frac{\partial}{\partial x} \hat{i} + \frac{\partial}{\partial y} \hat{j} + \frac{\partial}{\partial z} \hat{k}$$

Del
operator

apply to a
scalar

$$\nabla p = \frac{\partial p}{\partial x} \hat{i} + \frac{\partial p}{\partial y} \hat{j} + \frac{\partial p}{\partial z} \hat{k}$$

p pressure

$$\nabla P = \underbrace{\nabla_H P}_{\text{horizontal pressure gradient}} + \frac{\partial P}{\partial z} \hat{k}$$

horizontal
pressure
gradient

$$\nabla_H = \frac{\partial}{\partial x} \hat{i} + \frac{\partial}{\partial y} \hat{j}$$

points from low
values to
high values

$\vec{\nabla}$ apply to a vector

$$\nabla \cdot \vec{v} = \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} + \frac{\partial w}{\partial z}$$

$$\vec{v} = u\hat{i} + v\hat{j} + w\hat{k}$$

$$\frac{\partial}{\partial x}\hat{i} + \frac{\partial}{\partial y}\hat{j} + \frac{\partial}{\partial z}\hat{k}$$

$$\nabla \cdot \vec{v} = \nabla_{\hat{i}} \cdot v\hat{i} + \frac{\partial w}{\partial z}$$

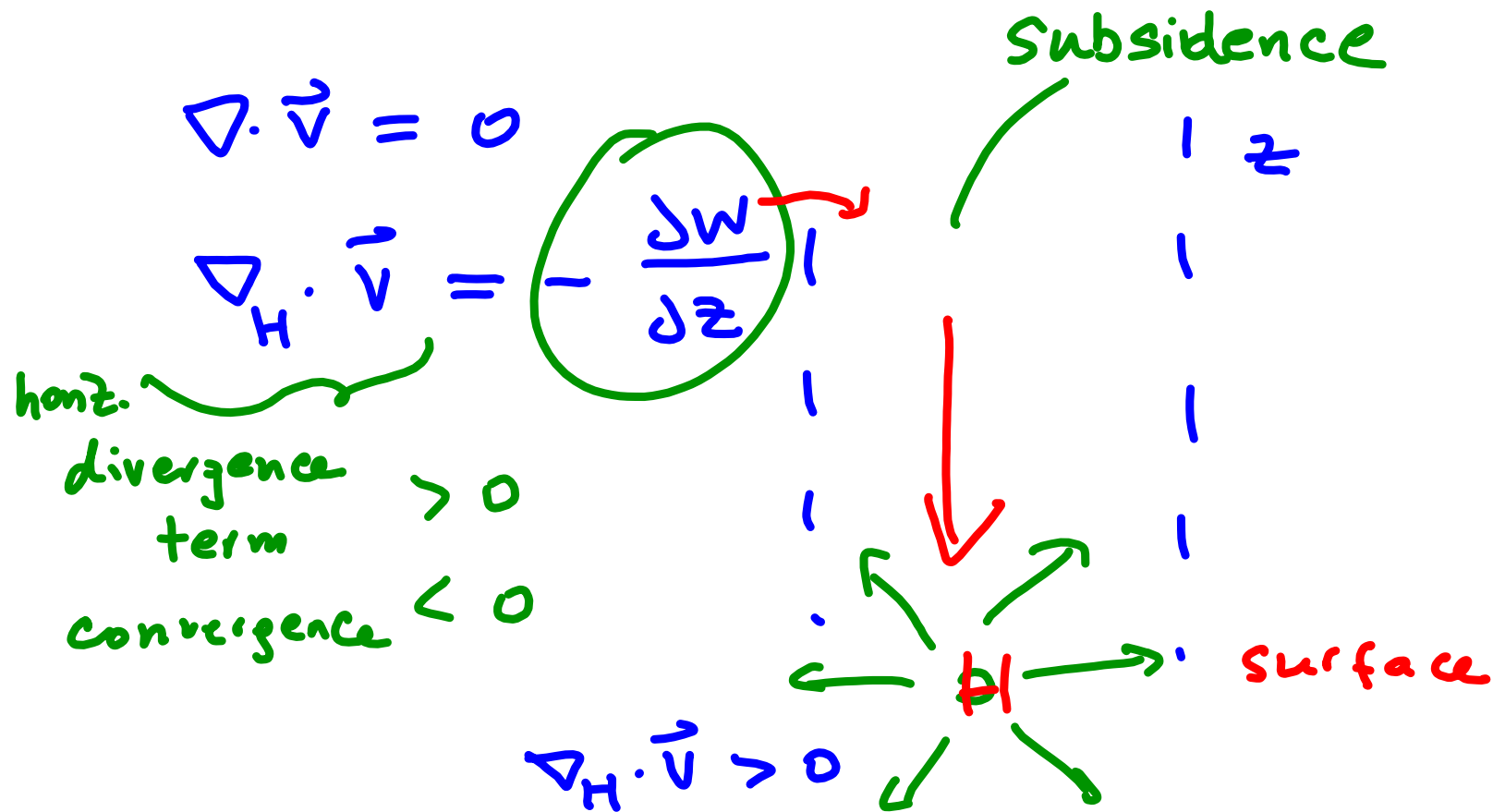
$$\nabla \cdot \vec{V} = \nabla_H \cdot \vec{V} + \frac{dW}{dz}$$

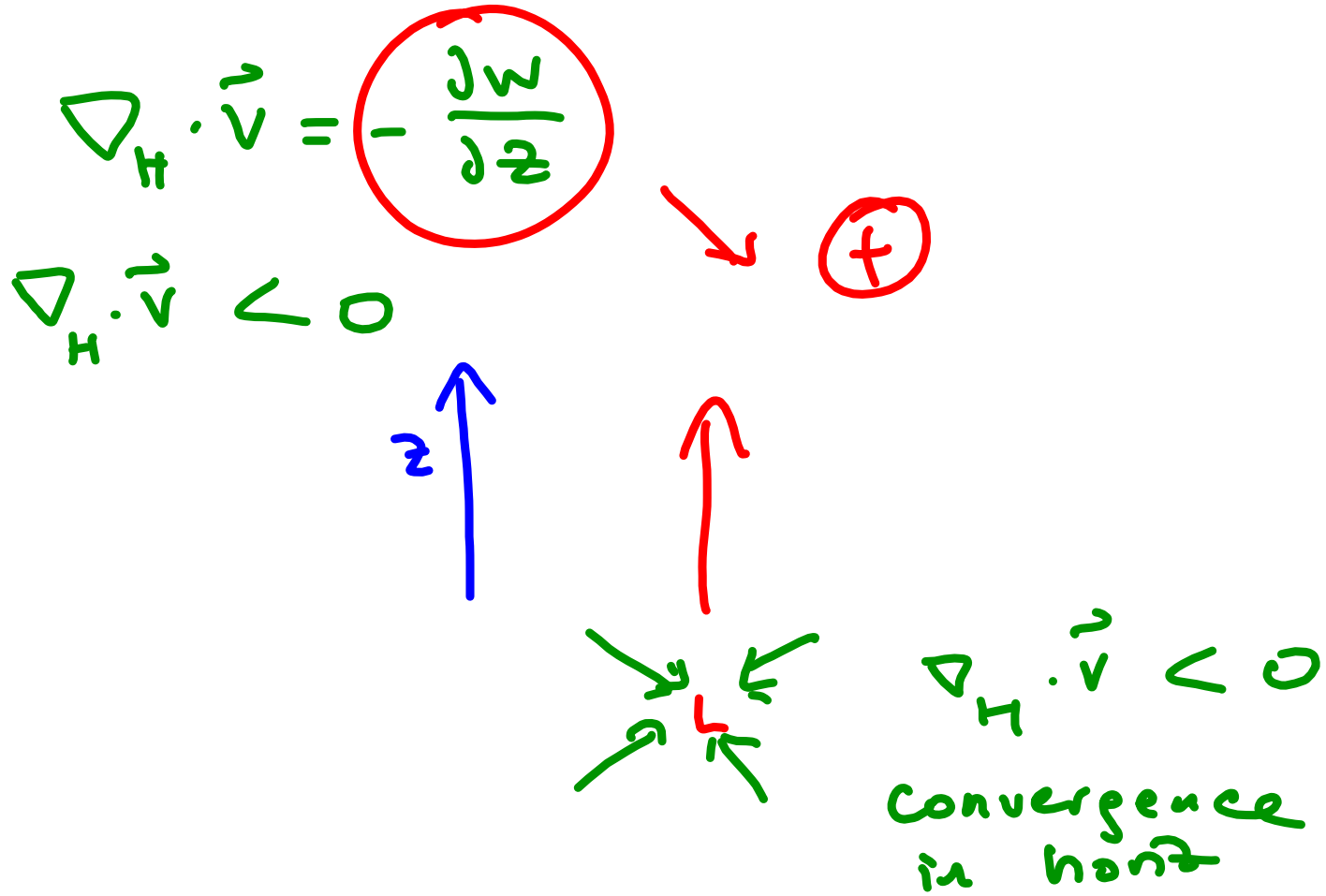
Application through mass continuity equation

* approximate / assume that the fluid is incompressible : $d\rho/dt = 0$ ρ density

Then mass continuity equation :

$$\nabla \cdot \vec{V} = 0$$





*

$$\frac{dT}{dt} = \frac{\partial T}{\partial t} + u \frac{\partial T}{\partial x} + v \frac{\partial T}{\partial y} + w \frac{\partial T}{\partial z}$$
$$= \frac{\partial T}{\partial t} + \vec{v} \cdot \nabla T$$

local change of Temp. in time

advection

Vorticity :

$$\nabla \times \vec{v} = \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ \partial/\partial x & \partial/\partial y & \partial/\partial z \\ u & v & w \end{vmatrix}$$

$\nabla \cdot \vec{v}$ divergence /
convergence

Back to governing equations

Atmosphere is a fluid - its motion is governed by the laws of physics such as Newton's 2nd law

$$m\vec{a} = \sum \vec{F}$$

$$m \frac{d\vec{v}}{dt} = \sum \vec{F}$$

not true
for Earth

valid in inertial
(non-accelerating)

frame of
reference

$$\vec{F}_{PGF} = -\frac{1}{\rho} \nabla p \quad \rho \text{ density}$$


Pressure Gradient Force

How do we know that?

- infinitesimal parcel - drop m term
 \uparrow mass

$$\frac{d\vec{v}}{dt} = \sum \vec{F} \quad \text{governing eqns}$$

1-3

$$\frac{dT}{dt} = \text{something}$$

cons. of mass

$$\frac{d\rho}{dt} = \text{other something}$$

Continuity eqn

$$\nabla \cdot \vec{v} = 0$$

in
incompr. fluid

Define fields that all depend
on x, y, z, t

$U(x, y, z, t)$

$V(x, y, z, t)$

$W(x, y, z, t)$

$P(x, y, z, t)$

$T(x, y, z, t)$

$q(x, y, z, t)$

$\rho(x, y, z, t)$

Zonal wind (x)

meridional wind (y)

vertical wind (z)

pressure

Temperature

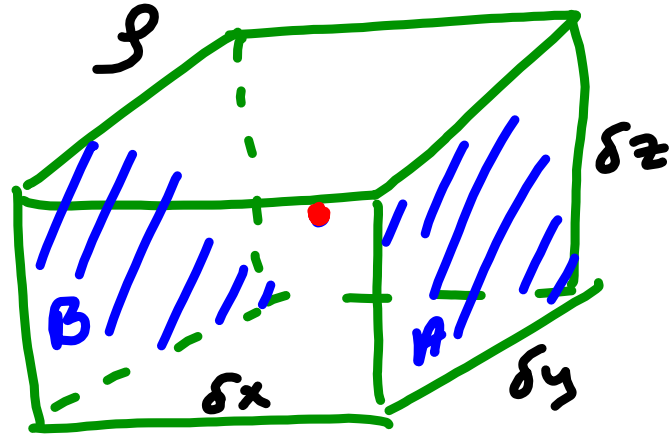
humidity

density

Pressure Gradient Force

- look at infinitesimal fluid element

Searching for
pressure exerted
on sides A & B
by the
atmosphere



put coordinate
system in the
middle

$$V = \delta x \delta y \delta z$$

$$m = \rho \delta x \delta y \delta z$$

$$\rho \cdot V$$

$$P_A = \frac{F_A}{A_A}$$

$$P_B = \frac{F_B}{A_B}$$

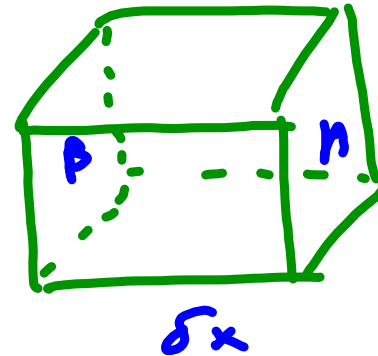
- pressure in the center p_0

Assume that pressure is continuous \Rightarrow
Taylor series

$$f(x) = f(x_0) + f'(x_0)(x - x_0) + \dots$$

$$p_A = p_0 + p' \frac{\delta x}{2} = p_0 + \frac{\partial p}{\partial x} \frac{\delta x}{2}$$

$$p_B = p_0 + p' \left(-\frac{\delta x}{2}\right) = p_0 - \frac{\partial p}{\partial x} \frac{\delta x}{2}$$



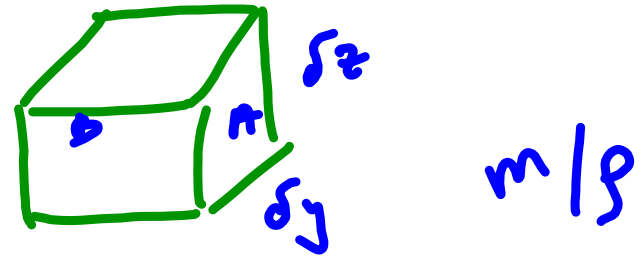
$$F_{A_x} = - \left(\cancel{p_0} + \frac{\partial p}{\partial x} \frac{\delta x}{2} \right) \delta y \delta z$$

$$+ F_{B_x} = + \left(\cancel{p_0} - \frac{\partial p}{\partial x} \frac{\delta x}{2} \right) \delta y \delta z$$

F_{A_x} is pointing in negative direction

$F = p \cdot A$

F_{B_x} is positive!



$$F_x = F_{A_x} + F_{B_x}$$

$$F_x = - \frac{\partial p}{\partial x} \delta x \delta y \delta z$$

$$F_x = \frac{m}{\rho} \left(-\frac{\partial p}{\partial x} \right)$$

$$\frac{F_x}{\cancel{m}} = -\frac{1}{\rho} \frac{\partial p}{\partial x}$$

neglecting
m

$$\left. \begin{aligned} F_x &= -\frac{1}{\rho} \frac{\partial p}{\partial x} \\ F_y &= -\frac{1}{\rho} \frac{\partial p}{\partial y} \\ F_z &= -\frac{1}{\rho} \frac{\partial p}{\partial z} \end{aligned} \right\}$$

$$\vec{F}_{PGF} = -\frac{1}{\rho} \nabla p$$

Fundamental forces

PGF

$$-\frac{1}{\rho} \nabla p$$

Gravitational
force

$$-g \hat{k}$$

Frictional
force

$$\vec{\tau}$$

friction

Apparent forces

Coriolis

Centrifugal

Coriolis force

$$f = 2\Omega \sin \phi$$

Coriolis
param.

$$-2\vec{\Omega} \times \vec{v} = -2 \begin{vmatrix} \hat{i} & \hat{j} & \hat{k} \\ 0 & -2\Omega \cos \phi & -2\Omega \sin \phi \\ u & v & w \end{vmatrix}$$

 Ω rotation
vector ϕ latitude

x comp.

$$2\Omega \sin \phi v - 2\Omega \cos \phi w$$

y

$$-2\Omega \sin \phi u$$

z

$$2\Omega \cos \phi u$$