

**1 On the Relationship Between Continuing Current**  
**2 and Positive Leader Growth**

Jeff L. Lapierre,<sup>1</sup> Richard G. Sonnenfeld,<sup>1</sup> Harald E. Edens,<sup>1</sup> Mike Stock,<sup>1</sup>

---

Corresponding author: Jeff Lapierre, Department of Physics, New Mexico Institute of Mining and Technology, Socorro, New Mexico, USA. (jlapierr@nmt.edu)

<sup>1</sup>Department of Physics and Langmuir  
Laboratory, New Mexico Institute of Mining  
and Technology, Socorro, New Mexico, USA

3 **Abstract.** It has long been speculated that the source of continuing cur-  
4 rent (CC) for a negative cloud-to-ground flash is provided by the growth of  
5 its positive leader into negative charge regions. In this study, data from the  
6 Langmuir Electric Field Array (LEFA) and Lightning Mapping Array (LMA)  
7 are used to investigate these speculations. LEFA and LMA data provide a  
8 way to estimate the occurrence and duration of CC and channel growth through-  
9 out a flash, respectively. By connecting LMA VHF sources onto contiguous  
10 channels, the growth of the positive leader associated with each return stroke  
11 is inferred. A linear correlation between positive-channel growth and CC du-  
12 ration is found, providing evidence that the positive leader grows with a con-  
13 stant velocity, but no obvious correlation of this velocity with CC occurrence  
14 is found. Each return stroke is then sorted by its channel growth rate and  
15 further identified by its CC type. This analysis also provides no identifiable  
16 correlation linking the positive-channel growth rate to CC occurrence or du-  
17 ration. Finally, the positive-channel-growth rate for the whole flash is cal-  
18 culated in 10-ms windows so that any trends occurring before, during, or af-  
19 ter the CC can be observed. This analysis too shows no correlation, which  
20 implies that positive-channel growth is not the primary mechanism that de-  
21 termines CC occurrence and duration.

## 1. Introduction

### 1.1. Continuing Current

22 Cloud-to-ground (CG) flashes consist of leaders that exit the parent cloud and connect  
23 to the ground. For a negative CG (-CG) flash, negative charge is carried to ground by a  
24 negative leader. Due to the bi-polar nature of lightning, the opposite end (referred to as  
25 the positive leader), typically located within the cloud, will be positively charged. When  
26 the negative leader connects to ground, it causes a surge of current called the return  
27 stroke, which may be followed by a steady, long-lived current called continuing current  
28 (CC). Why some return strokes are followed by CC and some are not is still not properly  
29 understood, but is thought to depend on the characteristics of the positive leader [*Krehbiel*  
30 *et al.*, 1979; *Rakov and Uman*, 1990; *Mazur*, 2002; *Saba et al.*, 2006a; *Williams*, 2006]. We  
31 combined VHF radio emissions measurements (that located regions of electric breakdown  
32 of air) with electric field change measurements (which observed charge motion) to study  
33 the relationship between the growth of these positive leaders and CCs.

### 1.2. Continuing Current and the Positive Leader

34 Among the first to propose a relationship between CC and the positive leader were  
35 *Krehbiel et al.* [1979], who suggested, “Because of the highly interactive nature of the dis-  
36 charge process it is likely that both effects (ie., channel negative resistance and availability  
37 of source current) are important factors in making the discharge discrete.” This statement  
38 is very general, however the phrase ‘availability of source current’ suggests some kind of  
39 leader growth in order to connect to additional sources of charge. *Rakov and Uman* [1990]  
40 supported this statement, “Given suitable conditions, the alternative between a continu-

ing current and a discrete stroke will then be a matter of the availability of a charge source capable of providing an appropriate current input to the channel.” Similarly, in *Rakov and Uman* [2003] on page 176, they stated, “The return stroke removes charge deposited on the channel by a preceding leader, whereas continuing current is likely to be associated with the tapping of fresh charge regions in the cloud”. More recently, *Saba et al.* [2006a] stated, “We can speculate that the higher occurrence of long continuing current is related to the availability of charges in the negative charge layer of the thunderstorm and thus to the horizontal extent of the thundercloud.” The phrase ‘in the negative charge layer’ implies that developing positive leaders are the charge source.

*Heckman* [1992] suggests a model which can be used to determine whether or not a channel can sustain CC. The two critical characteristics for this model are channel length and current on the channel. In summarizing the model developed in *Heckman* [1992], *Williams* [2006] states, “The extension of the channel into the electric field of space charge aloft provides for a quasi-steady current source” [*Williams*, 2006]. Therefore this model depends directly on the positive leader growth to provide sufficient current for CC to occur.

*Mazur* [2002] stated a direct link between CC and in-cloud leader activity, “The presence of continuing current, commonly observed either in the E-field change record, or as the continuing luminosity of a visible channel, is an indication of a developing leader in the flash. The duration of continuing current, which delineates the duration of leader development, varies from a few to hundreds of ms.” They go on to say, “The current source that maintains the arc is associated with the breakdown process at the leader tip and the self-propagation of the leader channel.”

Owing to these speculations, the analysis of the positive leader is paramount in understanding CC; the goal of this study.

### 1.3. Differing Regimes of Continuing Current

The definition of CC duration has been refined over the past several decades. *Brook et al.* [1962] and *Kitagawa et al.* [1962] defined long CC as a continuous electric field change lasting longer than 40 ms. *Shindo and Uman* [1989] then defined short CC as lasting between 10 and 40 ms. They also discussed what they called 'questionable' CC, which was CC that lasted less than 10 ms. However at the time it was difficult to determine if this questionable CC was actually due to CC, in-cloud events, or the tail of the return stroke electric field change. *Rakov and Uman* [2003] determined that the maximum duration of a return stroke was 3 ms, which *Ballarotti et al.* [2005] used to introduce very short CC. They defined very short CC as lasting between 3-10 ms. In this paper we will be using the definitions of long, short, and very short CC which are summarized in Table 1.

### 1.4. An Intuitive Model

The positive leader grows by the electric breakdown of air due to the large potential difference between the leader tip and the space charge in the cloud. As newly ionized channel grows in conductivity it increasingly approximates an equipotential. As this occurs charges must be redistributed. While we could not find data for positive channels, it has been found that newly formed negative channels have a line charge density between  $-0.02 \text{ mC m}^{-1}$  [*Warner et al.*, 2003] and  $-1.8 \text{ mC m}^{-1}$  [*Lu et al.*, 2011]. These results lead one to suppose that each time a positive leader extends, a similar amount of charge becomes available. Since it is assumed that the channel is grounded, there is a large

84 electric potential difference initially between the old channel and the new channel, which  
85 supplies current to the grounded channel as the electric potential equalizes. This picture  
86 suggests that current should be proportional to some power of the growth rate. If this is  
87 the case, then there should be some measurable increase in positive leader growth during  
88 CC compared to when there is none.

## 2. Instrumentation and Methods

89 This study focuses on  $-CG$  flashes that occurred around Langmuir Laboratory near  
90 Socorro, New Mexico. Data from the Langmuir Electric Field Array (LEFA, 0.3 Hz-50  
91 kHz electric field change)[*Sonnenfeld and Hager, 2013*] and Lightning Mapping Array  
92 (LMA, 60-66 MHz VHF band)[*Rison et al., 1999*] are used to analyze the dependence of  
93 CC on positive-channel branching and growth. The LEFA is an array of nine slow field  
94 change sensors sampling at 50 kHz. Figure 1 illustrates the locations of each station and  
95 Langmuir Laboratory, which is considered the origin for our coordinate system. A total of  
96 nine flashes, which occurred during the summers of 2012 and 2013, are analyzed. These  
97 include 57 return strokes, 30 of which are followed by CC (10 very short, 6 short, and 14  
98 long, per Table 1).

99 The LEFA data allow us to determine which return stroke was followed by CC and also  
100 the duration of the CC. In Figure 2, LEFA3 data for a bolt-from-the-blue flash occurring  
101 on 14 August 2012 is plotted, emphasizing the CC, which is represented by the black line  
102 segments. We calculate CC duration by measuring the time interval between the return  
103 stroke field change and its intersection with interstroke electric field activity (example  
104 shown in the inset of Figure 2). The intersection is decided by the measured electric field  
105 being within 0.1% of a linear fit to the interstroke electric field. Table 2 summarizes the

106 CC durations for the 14 August 2012 flash measured at six different LEFA stations. CC  
107 durations are shown in order of the distance between each station and the ground strike  
108 location (illustrated in the right plot of Figure 9 by a red diamond). Even though LEFA5  
109 is the closest station to the return stroke, we analyze LEFA3 data because of the electric  
110 field enhancement due to local terrain variations. LEFA3's location causes the electric  
111 field lines to converge and therefore generally gives it the greatest sensitivity of all the  
112 stations. Most of the stations are in agreement, however there are some discrepancies,  
113 which mainly have to do with the automated calculation of the duration. For example,  
114 for RS4, LEFA8, 10, 6, and 7 all measure the electric field waveforms that drop below the  
115 inter-stroke electric field before flattening out. This effect causes the automated estimation  
116 of the CC duration to be shorter than one would expect from inspection.

117 When measuring CC durations, high-speed video observations are preferable to E-field  
118 data, however, electric field measurements have been shown to also be very effective  
119 at observing CC. [Ross *et al.*, 2008]. Saba *et al.* [2006b] showed that the duration of  
120 CC determined from high-speed video ( $CC_{\text{video}}$ ) and from electric field measurements  
121 ( $CC_{\text{E-field}}$ ) agree. Figure 3 shows this relationship for 19 return strokes, with which  
122 they calculated an R-value of 0.87 between the two techniques. This result indicates  
123 that E-field measurements are an adequate proxy for estimating CC durations longer  
124 than approximately 100 ms. The data provided by Saba *et al.* [2006b] is lacking for CC  
125 durations less than 100 ms.

126 For electric field measurements performed near a lightning flash, it becomes difficult to  
127 distinguish between CC and in-cloud activity [Ross *et al.*, 2008]. This is solved by having  
128 multiple electric field measurements. The sign of the electric field change measured will

129 depend on the direction of motion of the in-cloud activity relative to the station. The  
130 same is not true for CC. Electric field changes for CC are well approximated as monopoles  
131 [*Krehbiel et al.*, 1979]. Therefore, if the sign of the electric field change for multiple stations  
132 agree, and it resembles what one would expect for CC, then the electric field change in  
133 question is due to CC. For example, compare the locations of LEFA3 and LEFA5 (see  
134 Figure 1) to the direction of the positive leader in the flash illustrated in Figure 2. The  
135 positive leader moves to the north west directly towards LEFA3. Using only LEFA3  
136 data, one would not be able to distinguish if the electric field changes occurring after the  
137 third, fourth and fifth return strokes were due to CC or the motion of the positive leader.  
138 However, since the positive leader motion is perpendicular to LEFA5 and it measures the  
139 same sign of electric field change (see Figure 4), one can be confident that CC is the cause  
140 of the electric field change. This analysis has been performed for all flashes in this study  
141 to legitimize that the electric field changes are due to CC.

142 LMA data were obtained from the Langmuir LMA comprising 28 stations situated  
143 around Langmuir Laboratory. The high number of instruments, along with the relatively  
144 quiet environment, gives us the sensitivity required to detect positive breakdown activity,  
145 which is much quieter in RF than that produced by the negative leader [*Thomas et al.*,  
146 2004; *Edens et al.*, 2012]. To determine channel growth, we modify the PulseGraph  
147 function described in *Hager et al.* [2007]. This function was designed to join LMA VHF  
148 source points provided to arrive at a channel structure for a complete flash which can  
149 be used to measure channel length. It can also be used to connect new points to an  
150 existing channel as the flash progresses in order to calculate the length increase of the  
151 channel during the time interval of interest. An example of this analysis is illustrated



152 in Figure 5. Although only the planar view is shown in Figure 5, the LMA provides  
153 three-dimensional locations of VHF source points; the channel lengths reported in this  
154 paper use three-dimensional data. The red lines on the LEFA and LMA data (left panels)  
155 highlight the time of the fifth return stroke, while the red line segments on the PulseGraph  
156 figure indicate the new channel growth during that CC. Figure 6 illustrates a plan view  
157 comparison between PulseGraph and LMA VHF sources. In order to limit the number of  
158 noise solutions and prevent false channel detection by the PulseGraph function, LMA data  
159 are filtered using relatively strict parameters of minimum number of stations and reduced  
160 chi squared values (typically around 12 and 2 respectively). Applying the PulseGraph  
161 function on the filtered LMA data correctly fits the branch structure found in the LMA  
162 data and yields the channel lengths used in this analysis. It should be noted that, while  
163 the filtering scheme does change the overall magnitude of the channel growth rates, it  
164 does not affect the characteristics described in this analysis.

### 3. Analysis and Discussion

165 Channel growth is determined for the nine flashes as described above. For each return  
166 stroke the time period of CC activity is determined from the LEFA waveform. The  
167 PulseGraph algorithm is used to calculate channel growth during that same time interval.  
168 The channel growths for all 57 return strokes are plotted in Figure 7 against CC duration.  
169 The left panel depicts the data as a base 10 log-log plot while the right panel is a linear  
170 plot of the same data. The primary reason for illustrating the data both ways is to  
171 show the highly linear dependence (with the log-log plot) while also clearly displaying the  
172 cumulative speed of the positive leader (with the linear plot). Calculating the slope of  
173 the log-log plot gives us a power law relation of 0.93 (i.e. growth of the positive leader

174 equals duration of CC to the power of 0.93). In other words, the relationship is almost  
 175 linear. At first glance the data appear to support the accepted hypothesis that CC is  
 176 caused by the growth of the positive leader because long CCs correspond to large channel  
 177 growths. However, since the exponent in the left panel of Figure 7 is roughly one, there is  
 178 a simpler explanation. If the average positive leader branch growth is constant (given by  
 179 the slope in the right panel of Figure 7), and continues throughout the flash, this linear  
 180 correlation would also occur. Observations presented later in this paper (seen in Figures  
 181 10, 11, and 12) show that the positive leader grows throughout the flash, whether or not  
 182 there is CC. Therefore, the apparent correlation between long CC and channel growth  
 183 appears because channel growth of long CCs was calculated for longer periods of time.

184 Due to this constant positive leader velocity, if the number of branches can be estimated,  
 185 then an estimate for the positive leader velocity can be found also. The branch analysis  
 186 begins after the final stepped leader occurs so as not to count branches arising as part of  
 187 the negative leader. From this time zero, the flash is considered in 100 ms increments. For  
 188 each increment a branch is considered active if it contains new VHF sources during this  
 189 time period. The number of branches per time window is averaged over the whole flash,  
 190 and these are averaged over all the nine flashes in this study. The results obtained from  
 191 this analysis show that positive leaders have on average approximately 10 active branches.

The slope found in Figure 7 represents the cumulative speed of all the active branches  
 in the positive leader. By dividing this slope by the estimate of the average number of  
 branches in the positive leader found above, an estimate for the velocity of an individual  
 positive leader branch can be found. Therefore,

$$v_{\text{pos}} = \frac{R_{\text{tot}}}{N_{\text{b}}} = \frac{2.1 \times 10^5 \text{ ms}^{-1}}{10} = 2.1 \times 10^4 \text{ ms}^{-1}, \quad (1)$$

192 where  $v_{\text{pos}}$  is the velocity of the positive leader,  $R_{\text{tot}}$  is the growth rate of the positive  
 193 leader (the slope of the waveform(s) in Figure 7), and  $N_b$  is the estimated average number  
 194 of branches of the positive leader. The value of  $2 \times 10^4 \text{ ms}^{-1}$  calculated agrees with what  
 195 is found in the literature. *Edens et al.* [2012] observed velocities of  $1 - 3 \times 10^4 \text{ ms}^{-1}$   
 196 in a triggered flash in New Mexico while *van der Velde and Montanyà* [2013] observed  
 197 horizontal velocities of  $1.6 - 2.6 \times 10^4 \text{ ms}^{-1}$  for natural flashes observed in Spain. However,  
 198 three-dimensional velocities have been observed as high as  $3.3 \times 10^6 \text{ ms}^{-1}$  [*Yoshida et al.*,  
 199 2010].

The data in Figure 7 do not indicate an increase in growth rate during CC, but there was much scatter in the data. Therefore a different approach is taken to find a relationship between growth rate and CC. We explicitly calculate an average growth rate ( $R_{\text{RS}}$ ) for each of the 30 return strokes using

$$R_{\text{RS}} = \frac{G}{t_{\text{CC}}}, \quad (2)$$

200 where  $G$  is the channel growth during the period of return stroke and CC, and  $t_{\text{CC}}$  is the  
 201 return stroke and CC duration. Organizing these results into a histogram gives Figure  
 202 8. The colors in Figure 8 represent return strokes with long (red), short (green), very  
 203 short (yellow), or without CC (blue) (according to the definitions of Table 1). Figure 8  
 204 is in support of what Figure 7 implied. For the lowest bin ( $0-250 \text{ km s}^{-1}$ ), we note that  
 205 17 of 29 (59 %) return strokes are followed by CC. Were the prior assumptions about  
 206 the relation between channel growth and CC valid (per our discussion in Section 1.2), we  
 207 would assume that the next bin would have a larger fraction of CC strokes, but in fact  
 208 only 8 of 19 flashes (42 %) are followed by CC. In the three bins with highest growth  
 209 rates, we have a paucity of data, but there is no evidence that the highest rates provide a

210 larger fraction of CC flashes. Taken as a whole, Figure 8 does not support the assumption  
211 that longer CCs correspond with higher positive channel growth rates.

212 In order to make the best use of our nine flashes, we applied still another form of data  
213 analysis. So far, the growth rates were averaged over the time frame of the return stroke  
214 and CC. However it is possible to determine the growth rate at higher time resolution. By  
215 dividing LMA data for an entire flash into 10 ms windows, the growth rate can then be  
216 calculated during those windows. These results are shown in Figures 10, 11 and 12 with  
217 the 10 ms growth plotted above the LEFA and LMA data for each flash. This analysis  
218 can identify possible trends in growth rate throughout the flash and help to determine  
219 if there is a correlation between positive-channel growth and CC not observable in the  
220 averaged growth rates from Figure 8.

221 Figures 6 and 10 illustrate a  $-CG$  flash occurring on 8 July 2013. This flash has  
222 seven return strokes according to National Lightning Detection Network (NLDN) data  
223 (highlighted by vertical blue lines and sequentially numbered in Figure 10), with the  
224 first and third return strokes connecting to ground at different locations. The first two  
225 return stroke ground points are indicated in Figure 6 by a magenta triangle. The third  
226 and subsequent return strokes go to ground at a different location, represented by red  
227 diamond in Figure 6. The stepped leaders are apparent in the LMA plot in Figure 10 by  
228 the low altitude VHF sources. The channel growth algorithm detects these two stepped  
229 leaders as growth maxima, which stand out in the top panel. Inspection of the LMA data  
230 show that after the second stepped leader (approximately 1.08 seconds in Figure 10) the  
231 growth rate remains constant throughout the rest of the flash. This constant growth rate  
232 persists even though the sixth and seventh return strokes are followed by long CC. The

233 constancy of positive-channel growth rate of the 20:28 flash is typical of seven of the nine  
234 flashes in this study.

235 Only two flashes in this study have growth rates that are not constant during positive-  
236 channel growth. The first is shown on the left panel of Figure 9 and Figure 11 which  
237 occurred on 8 July 2013. This flash has nine return strokes according to NLDN data  
238 (highlighted by vertical blue lines and sequentially numbered in Figure 11). There are  
239 three different return stroke locations (shown on the left panel of Figure 9). The first two  
240 return strokes ground locations are depicted as magenta triangles and and the remaining  
241 return strokes by the red diamond. Just as in the 20:28 flash, after the final stepped  
242 leader connects to ground (during the third return stroke at 0.525 s) the growth rate  
243 decreases and remains relatively constant. Even though the sixth return stroke initiates  
244 long CC, the growth rate does not immediately increase. At approximately 0.9 seconds  
245 in the figure, about 0.06 seconds after the initiation of the sixth return stroke, the growth  
246 rate begins to increase and peaks at around 1 second. There is another peak occurring  
247 at around 1.1 seconds, between the seventh and eighth return stroke. Unlike the previous  
248 peak this one occurs even though the seventh return stroke is not followed by CC. Finally,  
249 the ninth and final return stroke is followed by short CC, but is not accompanied by any  
250 appreciable increase in positive-channel growth rate.

251 The second flash without a constant growth rate is the bolt-from-the-blue discussed  
252 earlier, occurring on 14 August 2012, and illustrated on the right panel of Figure 9 and  
253 Figure 12. This flash contains six return strokes according to NLDN data (highlighted  
254 by vertical blue lines and sequentially numbered in Figure 11), with the third and fourth  
255 being followed by short CC and the fifth by long CC. The data show that the growth rate

256 increases just after the second return stroke, which occurs at approximately 0.5 seconds  
257 or midway through the time interval colored green. This point in time coincides with  
258 the time when profuse branching of the positive leader begins (compare Figure 12 with  
259 the right panel of Figure 9). This coincidence indicates that there may be a correlation  
260 between growth rate and CC. However later in the flash, just before the return stroke  
261 with long CC begins, there is a drop in growth rate, which coincides with a decrease in  
262 branching. The fact that the growth rate decreases during the longest CC supports our  
263 earlier observation that CC duration does not depend in any simple way on the growth  
264 rate of the positive leader.

265 In summary, seven of nine flashes showed no significant variation in positive-channel  
266 growth rate, despite the instance of CC. Even though two of nine flashes do have non  
267 constant growth rates, these showed trends that were inconsistent with the speculation  
268 that the growth of the positive leader is the source of CC. While some increases in growth  
269 rates did coincide with the occurrence of CC, others did not. Also, there were instances  
270 where long CC was accompanied by lower growth rates than shorter duration CC.

#### 4. Conclusions

271 Using LEFA data to find CC duration and the PulseGraph function to estimate chan-  
272 nel growth vs time, the growth of positive leaders was compared to CC occurrence and  
273 duration. Plotting growth vs CC duration for individual return strokes indicates that the  
274 positive leader grows at a constant rate. Using the growth rate, obtained from the linear  
275 slope of this plot, we were able to estimate the velocity of the positive leader, which agrees  
276 with previous values reported in the literature.

277 The average growth rate during each return stroke, including CC duration, was calcu-  
278 lated. Each return stroke was then categorized by their average growth rate and compared  
279 by their CC type (long, short, very short, or no CC). We found no significant difference  
280 in growth rate based on CC type.

281 The growth rate throughout a flash during 10 ms windows was analyzed for the nine  
282 flashes in this study. Seven out of the nine flashes contained constant growth rates during  
283 positive leader activity even though there were occurrences of CC, while the remaining  
284 two flashes showed peaks which are inconsistent with CC occurrence.

285 It is possible that CC is caused by growth that occurs on a small scale unresolved  
286 by the LMA. However, the observations in this study all agree with the following: the  
287 growth of the positive leader is not the primary mechanism determining CC occurrence and  
288 duration. Therefore, there must be some other mechanism that determines the occurrence  
289 and duration of CC.

290 Most of the recent papers concerning continuing current focus on observations and  
291 assume that the mechanism is known. As mentioned in Section 1.2, many researchers  
292 believe that the source of CC comes from the growth of the positive leader, therefore  
293 studies into other mechanisms as the source of CC are lacking. *Heckman* [1992] does a  
294 thorough analysis of why some lightning flashes get cut-off and produce multiple return  
295 strokes and some produce CC, or both. However, he also assumes that the current source  
296 is provided by the positive leader growth. Similarly, *Mazur* [2002] models how the bottom  
297 layers of branching can screen upper branches from the ground electric field and result  
298 in the cutoff of the channel to ground, but again assumes that the source of current

299 comes from the channel growth. Our current findings indicate that more thought must  
300 be dedicated to alternate mechanisms of continuing current.

301 **Acknowledgments.** Raw data and scripts to reproduce the figures presented in this  
302 paper may be found at [www.physics.nmt.edu/~rsonnenf/papers/2014CCandLEADERS/](http://www.physics.nmt.edu/~rsonnenf/papers/2014CCandLEADERS/)  
303 This work was funded by grant CMB-0724750 from the National Science Foundation.  
304 Additional funding was obtained from the DARPA-NIMBUS program. We gratefully  
305 acknowledge Vaisala Corporation for providing access to precision National Lightning  
306 Detection Network data. We thank the National Forest Service for the use of land at  
307 Langmuir Laboratory and the Bureau of Land Management for allowing siting of LEFA  
308 stations on Federal land.

## References

- 309 Ballarotti, M. G., M. M. F. Saba, and O. P. Jr, High-speed camera observations of  
310 negative ground flashes on a millisecond-scale, *Geophys. Res. Lett.*, *32*, L23,802, doi:  
311 10.1029/2005GL023889, 2005.
- 312 Brook, M., N. Kitagawa, and E. J. Workman, Quantitative study of strokes and continuing  
313 currents in lightning discharges to ground, *J. Geophys. Res.*, *67*, 649–657, 1962.
- 314 Edens, H. E., et al., VHF lightning mapping observations of a triggered lightning flash,  
315 *Geophys. Res. Lett.*, *37*, L19,807, doi:10.1029/2012GL053666, 2012.
- 316 Hager, W. W., R. G. Sonnenfeld, B. C. Aslan, G. Lu, W. P. Winn, and W. L. Boeck, Anal-  
317 ysis of charge transport during lightning using balloon-borne electric field sensors and  
318 lightning mapping array, *J. Geophys. Res.*, *112*, D18,204, doi:10.1029/2006JD008187,  
319 2007.



- 320 Heckman, S. J., Why does a lightning flash have multiple strokes?, Ph.D. thesis, Mas-  
321 sachusetts Institute of Technology, 1992.
- 322 Kitagawa, N., M. Brook, and E. J. Workman, Continuing currents in cloud-to-ground  
323 lightning discharges, *J. Geophys. Res.*, *67*, 637–647, 1962.
- 324 Krehbiel, P. R., M. Brook, and R. A. McCrory, An analysis of the charge structure of  
325 lightning discharges to ground, *J. Geophys. Res.*, *84*(C5), 2432–2456, 1979.
- 326 Lu, G., W. Winn, and R. G. Sonnenfeld, Charge transfer during intracloud light-  
327 ning from a time-dependant multidipole model, *J. Geophys. Res.*, *117*, 16, doi:  
328 10.1029/2010JD014495, 2011.
- 329 Mazur, V., Physical processes during development of lightning flashes, *C. R. Physique*, *3*,  
330 1393–1409, 2002.
- 331 Rakov, V. A., and M. A. Uman, Long continuing current in negative lightning ground  
332 flashes, *J. Geophys. Res.*, *95*(D5), 5455–5470, 1990.
- 333 Rakov, V. A., and M. A. Uman, *Lightning Physics and Effects*, Cambridge University  
334 Press, Cambridge, 2003.
- 335 Rison, W., R. Thomas, P. Krehbiel, T. Hamlin, and J. Harlin, A gps-based three-  
336 dimensional lightning mapping system: Initial observations in central new mexico, *Geo-*  
337 *phys. Res. Lett.*, *26*(23), 3573–3576, 1999.
- 338 Ross, M., S. A. Cummer, T. K. Nielsen, and Y. Zhang, Simultaneous remote electric and  
339 magnetic field measurements of lightning continuing currents, *J. Geophys. Res.*, *113*,  
340 D03,101, doi:10.1029/2008JD010294, 2008.
- 341 Saba, M. M. F., M. G. Ballarotti, and O. P. Jr., Negative cloud-to-ground lightning  
342 properties from high speed video observations, *J. Geophys. Res.*, *111*, D03,101, doi:

- 343 10.1029/2005JD006415, 2006a.
- 344 Saba, M. M. F., O. P. Jr., and M. G. Ballarotti, Relationship between lightning return  
345 stroke peak current and following continuing current, *Geophys. Res. Lett.*, *33*, L23,807,  
346 doi:10.1029/2006GL027455, 2006b.
- 347 Shindo, T., and M. A. Uman, Continuing current in negative cloud-to-ground lightning,  
348 *J. Geophys. Res.*, *94 D4*, 5189–5198, 1989.
- 349 Sonnenfeld, R. G., and W. W. Hager, Electric field reversal in sprite electric field signature,  
350 *Mon. Weather Rev.*, *141*, 1731–1735, doi:10.1175/MWR-D-12-00220.1, 2013.
- 351 Thomas, R. J., P. R. Krehbiel, W. Rison, S. J. Hunyady, W. P. Winn, T. Hamlin, and  
352 J. Harlin, Accuracy of the Lightning Mapping Array, *J. Geophys. Res.*, *109*, D14,207,  
353 doi:10.1029/2004JD004549, 2004.
- 354 van der Velde, O. A., and J. Montanyà, Aircraft observations of a lightning channel in  
355 steps, *J. Geophys. Res.*, *118*, 13, 504–13, 519, doi:10.1002/2013JD020257, 2013.
- 356 Warner, T., J. H. H. Jr., and A. G. Detwiler, Aircraft observations of a lightning channel  
357 in steps, *Geophys. Res. Lett.*, *30*, 19, doi:10.1029/2003GL017334, 2003.
- 358 Williams, E. R., Problems in lightning physics - the role of polarity asymmetry, *Plasma*  
359 *Sources Science and Technology*, *15*, S91–S108, doi:10.1088/0963-0252/15/2/S12, 2006.
- 360 Yoshida, S., et al., Three-dimensional imaging of upward positive leaders in triggered  
361 lightning using vhf broadband digital interferometers, *Geophys. Res. Lett.*, *37*, L05,805,  
362 doi:10.1029/2009GL042065, 2010.

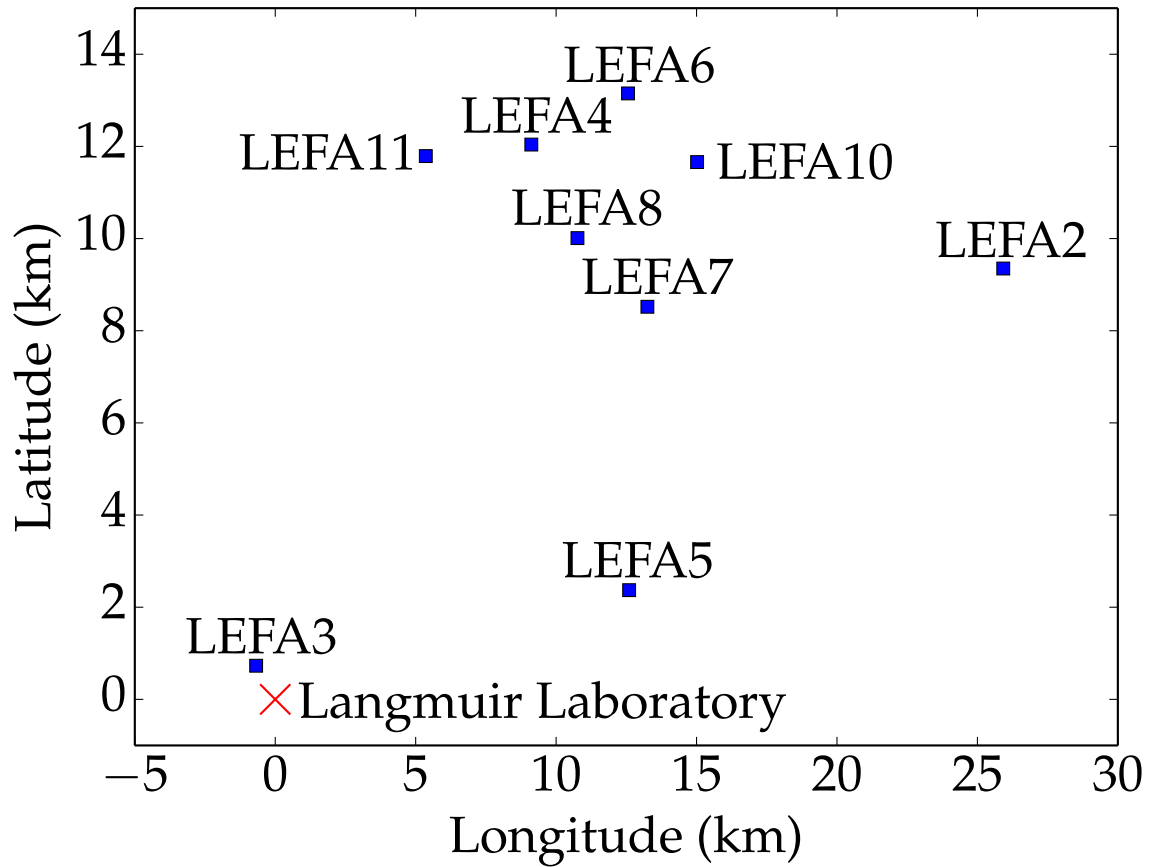
**Table 1.** Summary of CC definitions

Definition	Duration (ms)
Very Short	3-10
Short	10-40
Long	> 40

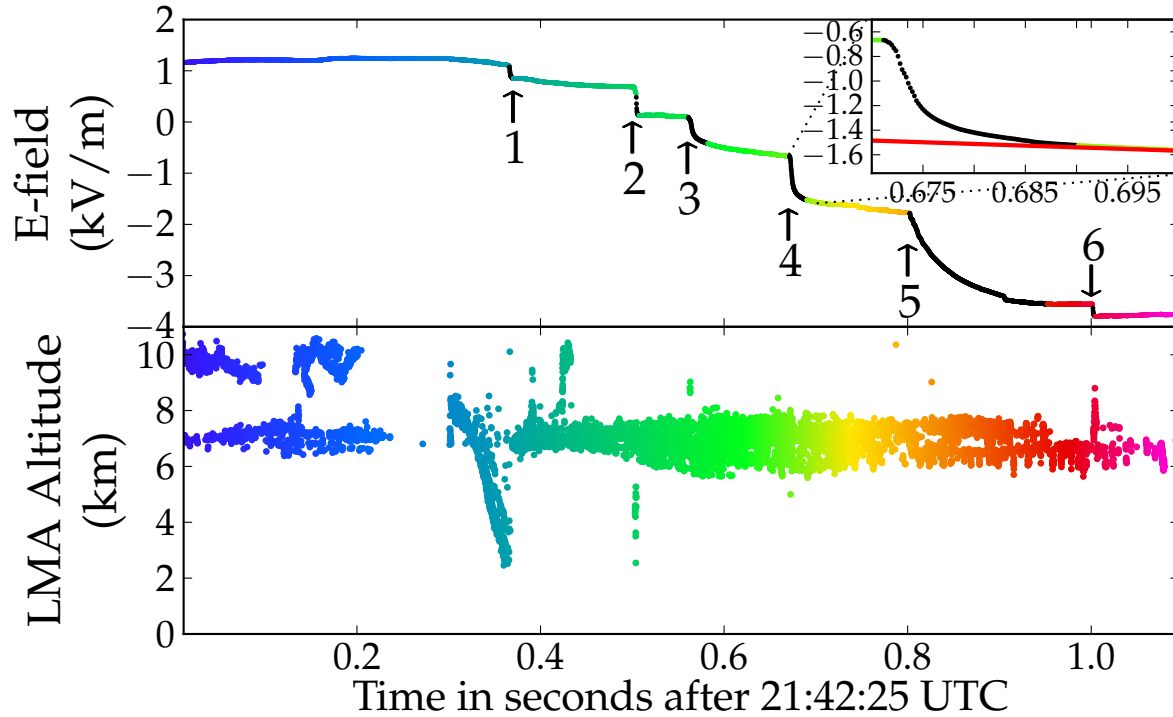
**Table 2.** Comparison of CC durations for 14 August 2012 flash as detected from multiple LEFA stations

Inst. #	RS1	RS2	RS3	RS4	RS5	RS6	RS Distance
	(ms)	(ms)	(ms)	(ms)	(ms)	(ms)	(km)
<b>LEFA5</b>	5.8	4.4	29.3	35.3	157	3.5	25.0
<b>LEFA3</b>	5.5	2.5	42.9	28.9	142	2.7	29.4
<b>LEFA7*</b>	10.0	4.3	31.0	22.2	156	3.5	30.9
<b>LEFA8</b>	5.0	5.0	33.9	23.6	158	2.6	32.8
<b>LEFA10</b>	6.7	4.1	32.7	19.2	159	3.5	33.8
<b>LEFA6*</b>	3.9	2.7	23.1	15.8	101	1.5	35.6

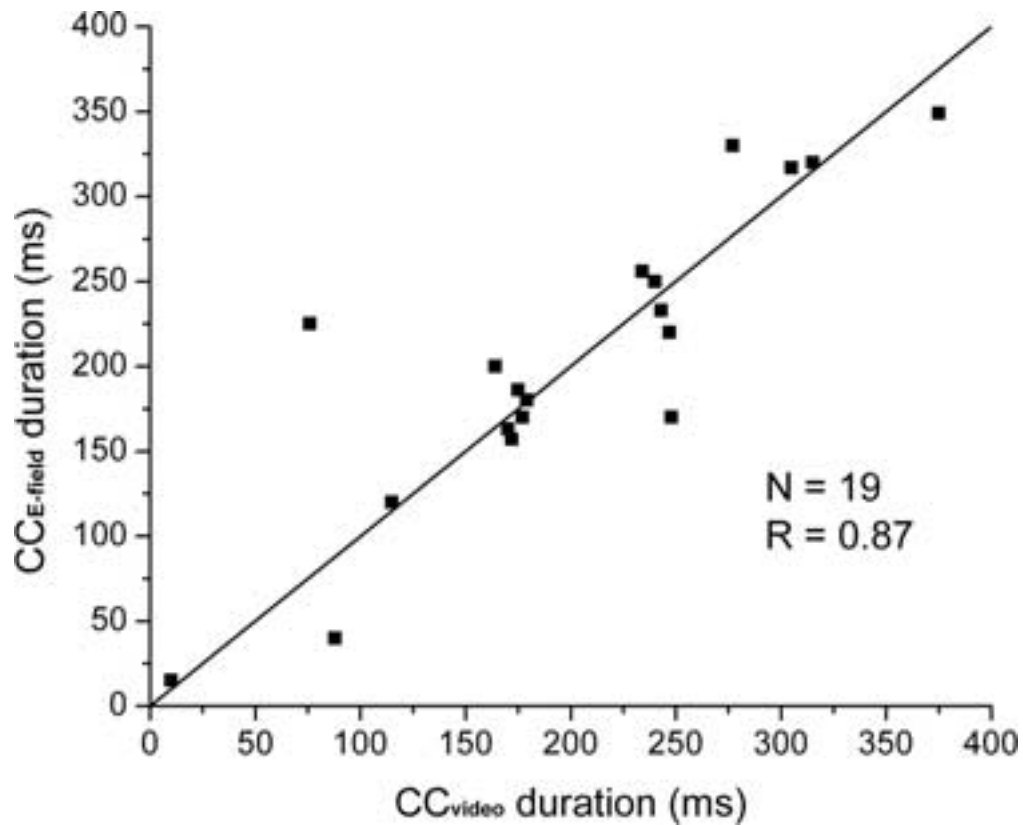
\*These instruments were noisy due to their proximity to power lines, and therefore provide much less reliable durations.



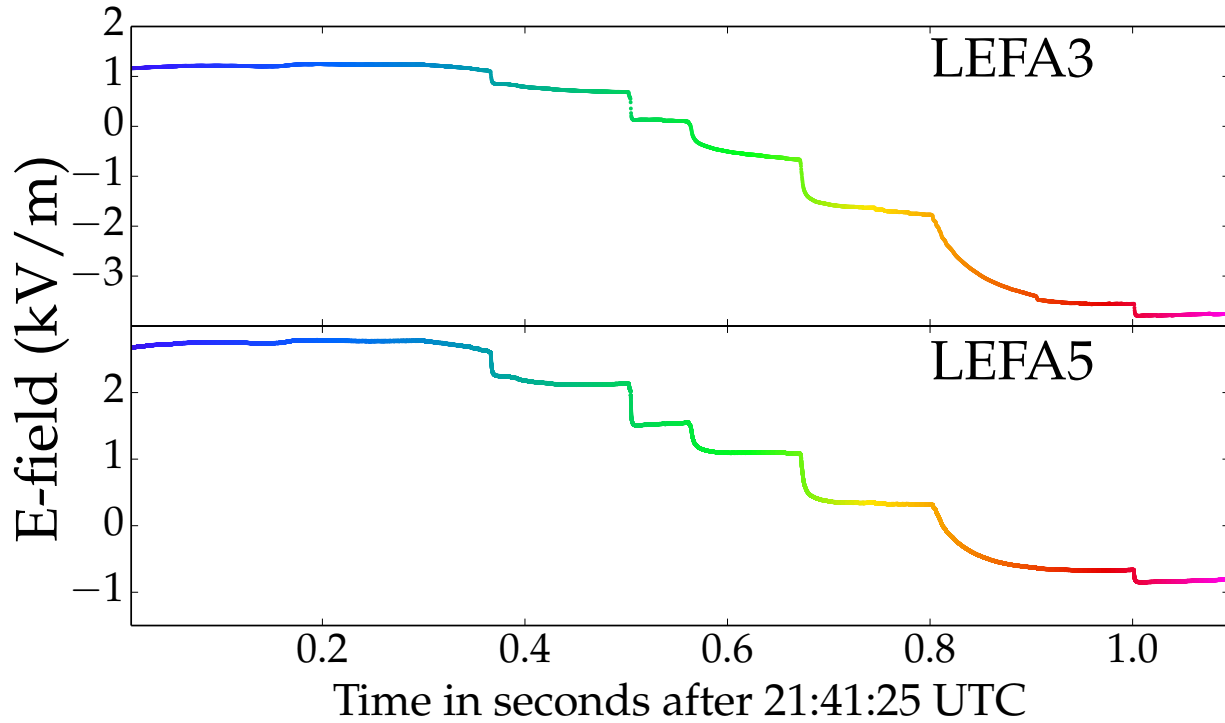
**Figure 1.** Figure illustrating the locations of the LEFA stations. Langmuir Laboratory is used as the origin for all the figures in this paper, which is labeled with a red X. LEFA2 is located on the campus of New Mexico Institute of Mining and Technology.



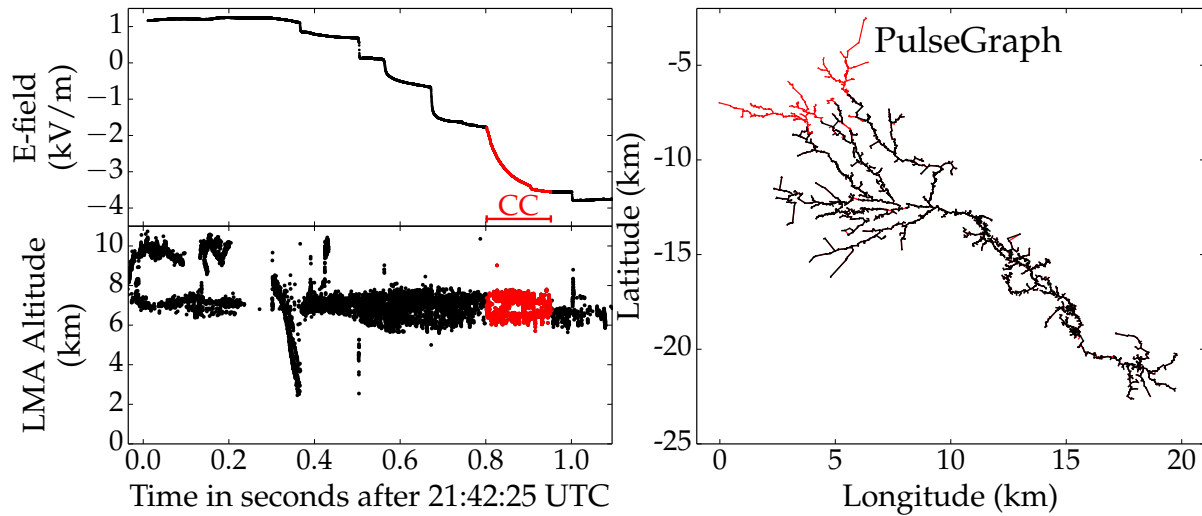
**Figure 2.** LEFA (top panel) and LMA (bottom panel) data for a bolt-from-the-blue flash occurring on 14 August 2012. This figure illustrates the method used to determine CC duration. The flash had six return strokes, three of which had CC (3, 4, and 5). The LMA data shows that most of the breakdown activity occurring after the first return stroke is in the negative charge region, which is caused predominately by positive leader breakdown. Color represents time from initiation of flash. The black line segments represent the duration of CC. The top inset demonstrates the technique used to identify the end of CC, which is determined by the point where the electric field and linear fit intersect.



**Figure 3.** Comparison between CC duration measured using electric field data and high-speed video data for 19 return strokes (figure from *Saba et al.* [2006b], used with permission). This result indicates that E-field measurements are an adequate proxy for estimating CC durations.

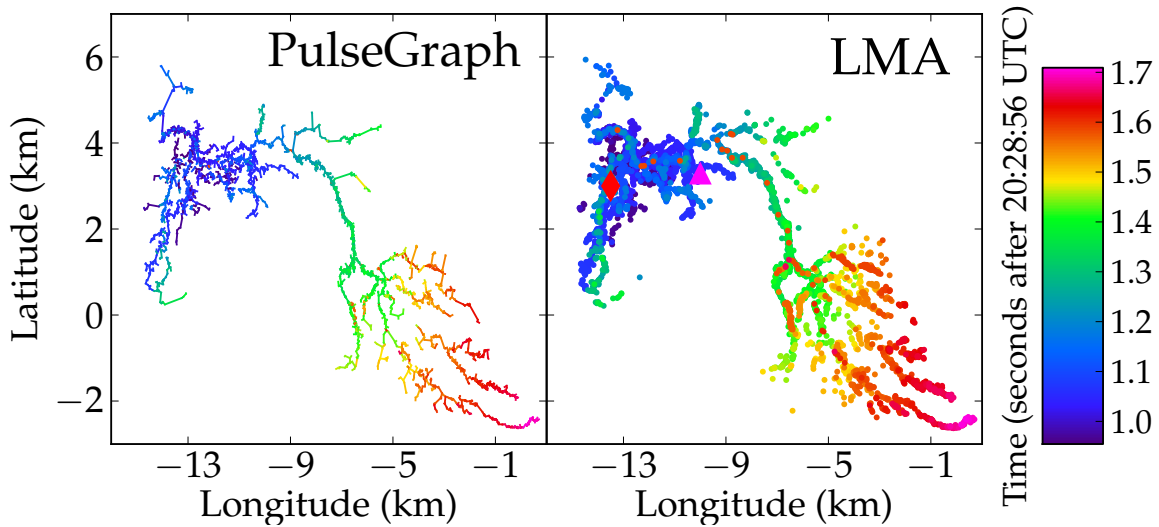


**Figure 4.** Electric field data from LEFA3 (top) and LEFA5 (bottom). The direction of the positive leader for this flash is moving at different directions compared to the location of each station (compare Figures 1 and 5). Since the sign of the electric field change is the same for each station, one can be confident that the electric field change is caused by CC and not the motion of the positive leader.

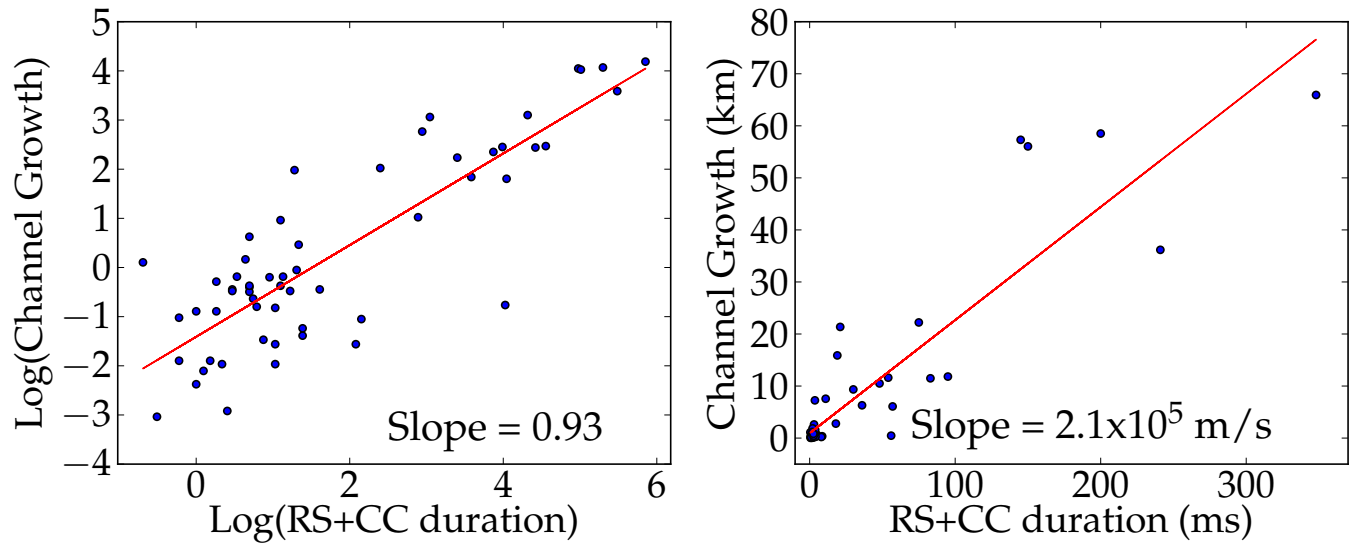


**Figure 5.** LEFA (top left), LMA VHF source altitude (bottom left), PulseGraph output (right) for bolt-from-the-blue flash occurring on 14 August 2012. This figure illustrates the method used to determine positive-channel growth during CC. Red on the PulseGraph plot indicates new channel growth during the time of the fifth return stroke (indicated also in red on the LEFA and LMA plots).

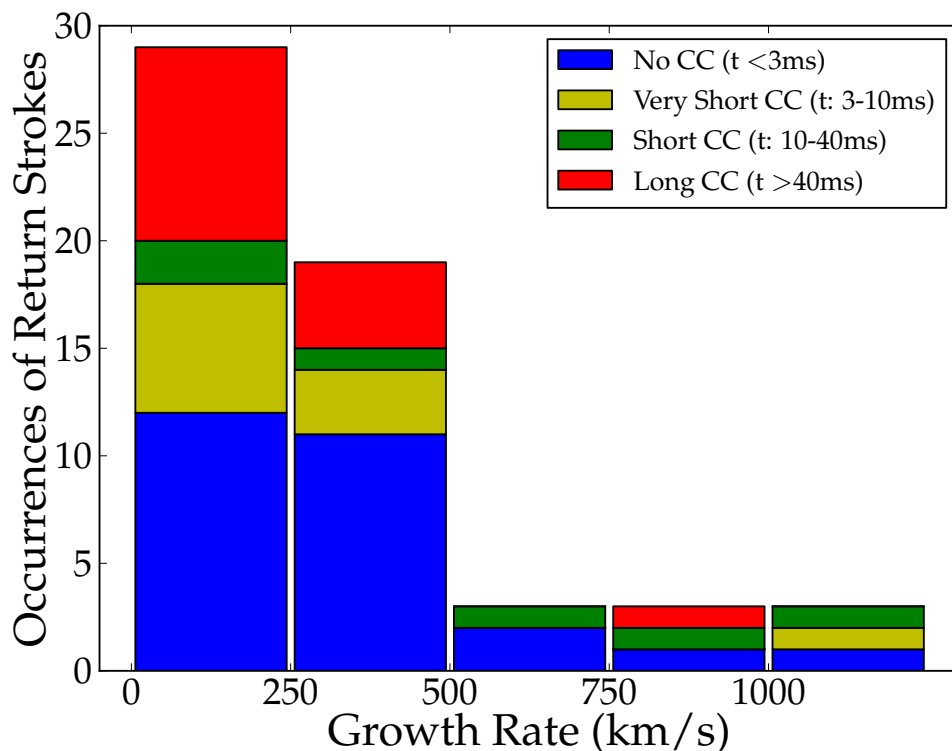




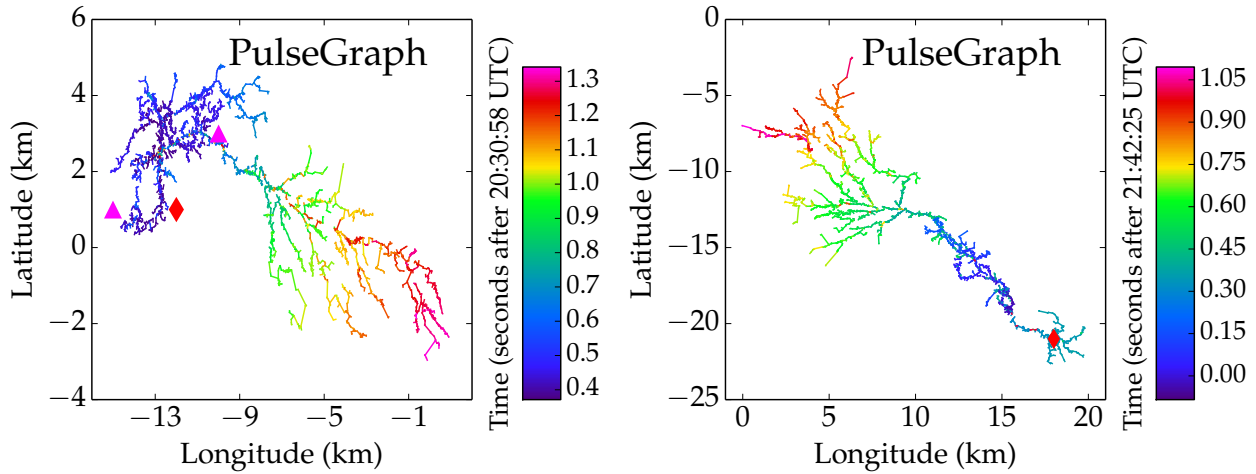
**Figure 6.** LMA (right) and PulseGraph (left) plan views of  $-CG$  flash occurring at 20:28:56 UTC on 8 July 2013. This figure demonstrates that the PulseGraph algorithm correctly fits the branch structure found in the LMA data. Color represents time (shown in colorbar). The magenta triangle designates the location of the first return stroke while the red diamond represents the location of the rest of the return strokes according to NLDN data.



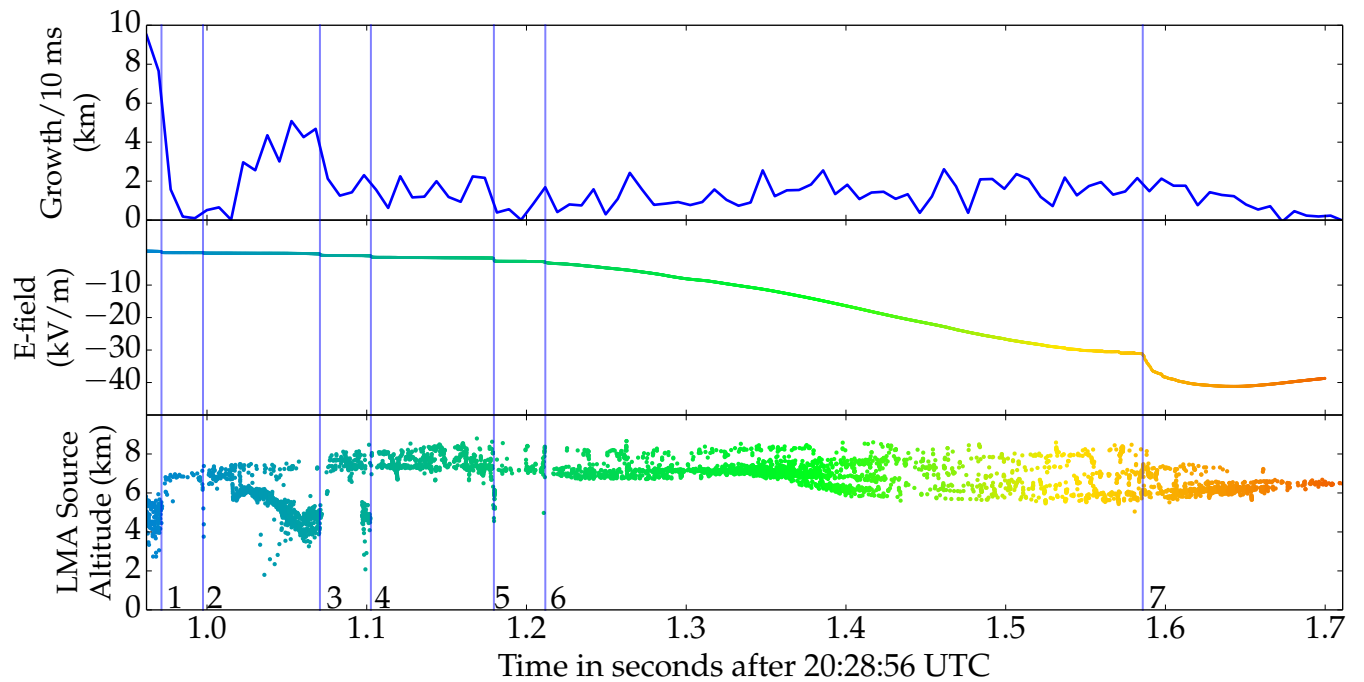
**Figure 7.** Plot of channel growth vs CC duration depicted as a base 10 log-log plot (left) and linear plot (right) for 9 flashes with a total of 57 return strokes. The log-log plot illustrates that the data follows a power law function, with the power being 0.93, or nearly linear. The slope in the linear plot represents the cumulative speed of all the branches in the positive leader. This cumulative speed is used to estimate the velocity of each individual positive leader branch.



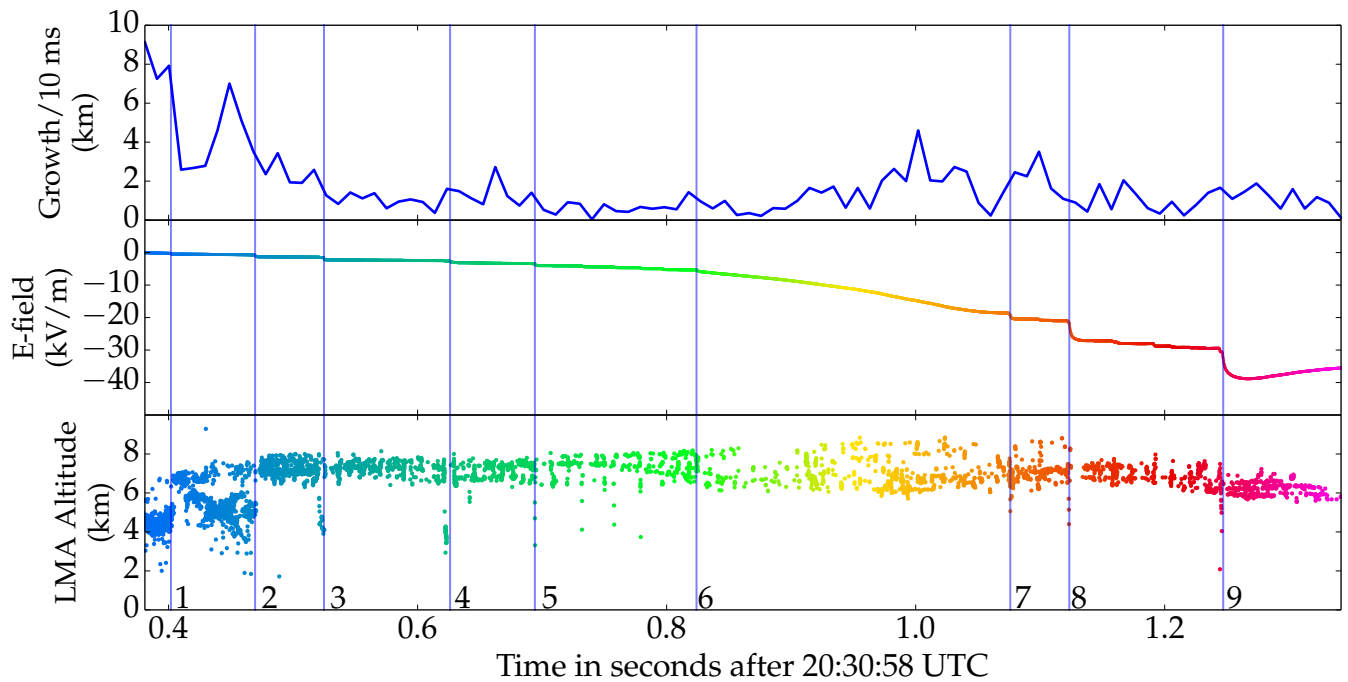
**Figure 8.** Histogram of return strokes categorized by growth rate. The y-axis shows the number of return strokes (from our data set of nine flashes) that had positive-channel growth rates indicated on the x-axis. The red, green, and yellow bars count long, short, and very short CC (as described in Table 1), while the blue bars represent no measurable CC (lasting less than 3 ms). The data reveals no preferential growth rate based on CC type.



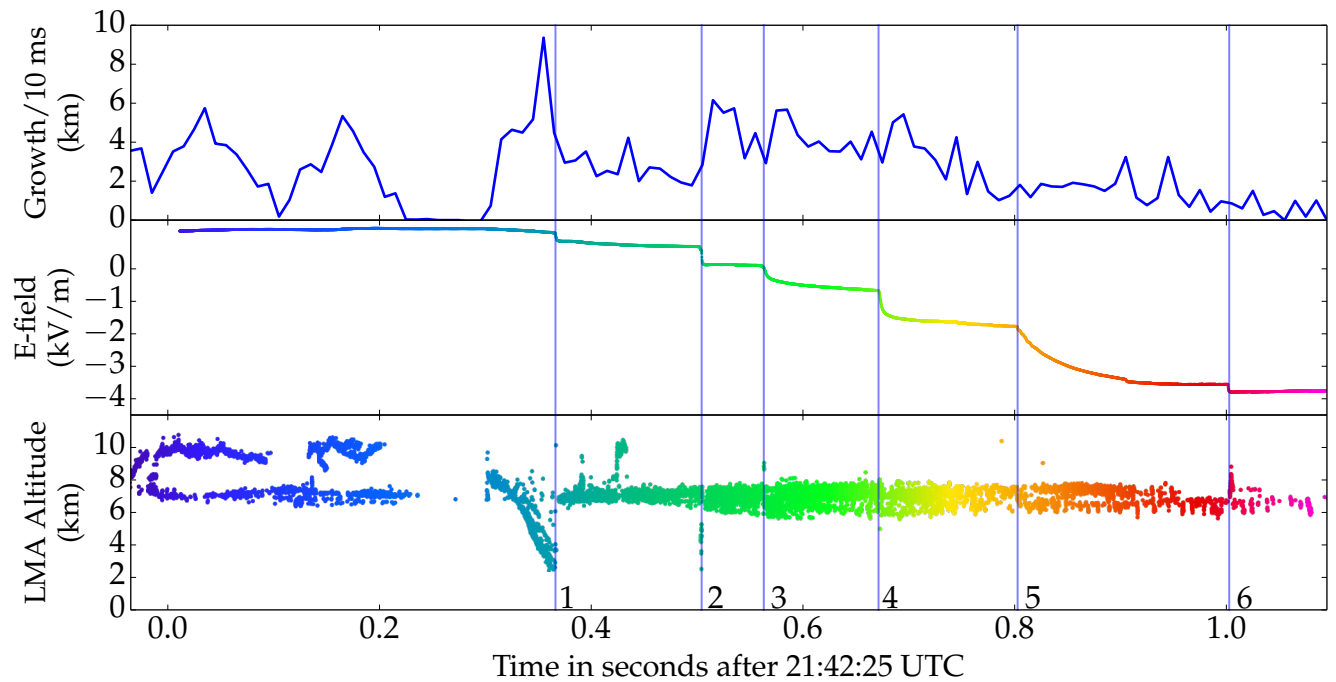
**Figure 9.** PulseGraph plan views of  $-CG$  flash occurring at 20:30:58 UTC on 8 July 2013 (left) and a bolt-from-the-blue flash occurring at 21:42:25 UTC on 14 August 2012 (right). Color represents time (shown in colorbars). The magenta triangles represent single return stroke locations, while red diamond designates the location of all the subsequent return strokes of that flash according to NLDN data. These PulseGraphs were used to calculate the growth rates for Figures 11 and 12.



**Figure 10.** Growth calculated over 10 ms windows throughout the flash (top). LEFA (middle) and LMA (bottom) data matched in time for a  $-CG$  flash occurring on 8 July 2013. This flash has seven return strokes (highlighted by vertical blue lines and sequentially numbered) according to NLDN data. Once the final negative stepped leader connects to ground, the positive-channel growth rate remains constant even though CC occurs.



**Figure 11.** Growth calculated over 10 ms windows throughout the flash (top). LEFA (middle) and LMA (bottom) data matched in time for a  $-CG$  flash occurring on 8 July 2013. This flash has nine return strokes (highlighted by vertical blue lines and sequentially numbered) according to NLDN data. Once the final negative stepped leader connects to ground, the positive-channel growth remains relatively constant until approximately 0.9 s. This is true even though the sixth return stroke is followed by long CC. Also, although the seventh return stroke does not initiate CC, there is a peak in growth rate between the seventh and eighth return stroke.



**Figure 12.** Growth calculated over 10 ms windows throughout the flash (top). LEFA (middle) and LMA (bottom) data matched in time for a bolt-from-the-blue flash occurring on 14 August 2012. This flash has six return strokes (highlighted by vertical blue lines and sequentially numbered) according to NLDN data. By comparing this data with Figure 9 it is apparent that the growth rate increases when the number of active branches increases.